

---

# Tracking down the elusive charginos / neutralinos through $\tau$ leptons at the Large Hadron Collider.

Amitava Datta

IISER, Kolkata, India

WORK DONE WITH IN COLLABORATION WITH

NABANITA BHATTACHARYYA (IISER KOLKATA)

Published in PRD 80,055016, (2009),hep-ph/0906.1460

# Introduction

---

Current lower limits on sparticle masses:

Sleptons ( $\tilde{l}$ ), lighter chargino ( $\tilde{\chi}_1^\pm$ )  $\sim 100 GeV$  (LEP)

Squarks ( $\tilde{q}$ )-Gluino ( $\tilde{g}$ )  $\sim 300 - 350 GeV$  (TEVATRON)

A **Model Independent** search strategy at LHC designed for sparticle masses just **above these limits** is called for.

These has not been done systematically.

**Discovery difficult if squarks and gluinos beyond LHC reach;**  
**sleptons and electroweak gauginos are little above the LEP limit.**

**Can we see the SUSY signal ?**

# Introduction

---

This scenario is consistent with all experimental data including DM (to be discussed later) but **it violates naturalness condition.**

This scenario is somewhat similar to the **Split SUSY** model.

# Introduction

---

Cannot exclude out this scenario since **SUSY breaking mechanism unknown**

**Essential reason:** chargino and second lightest neutralino decays to final states with lighter stau killing the clean trilepton signal, second lightest neutralino also decays invisibly making matters more complicated

A possible scenario **non-universal gaugino masses:**  
 $M_3 \gg M_2, M_1, m_0$  (**common slepton mass**) small  
(to be discussed later)

**Proposed Signal**

$2l + 1 \tau$  or  $1l + 2 \tau$

# Introduction

---

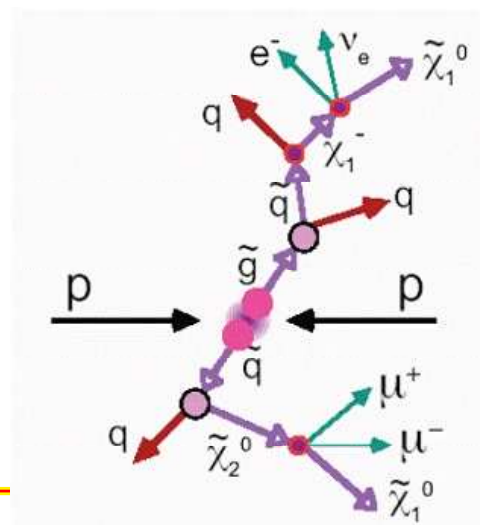
## Main Discovery Channels at the LHC:

### Squark -Gluino production followed by cascade decays

Generic SUSY Signals:

$m$  -Jets +  $n$  -leptons +  $\cancel{E}$  ; lepton = e,  $\mu$

Hardly identifies the underlying model. At least some of the sparticle masses should be reconstructed  $\rightarrow$  underlying SUSY breaking mechanism (with some luck).

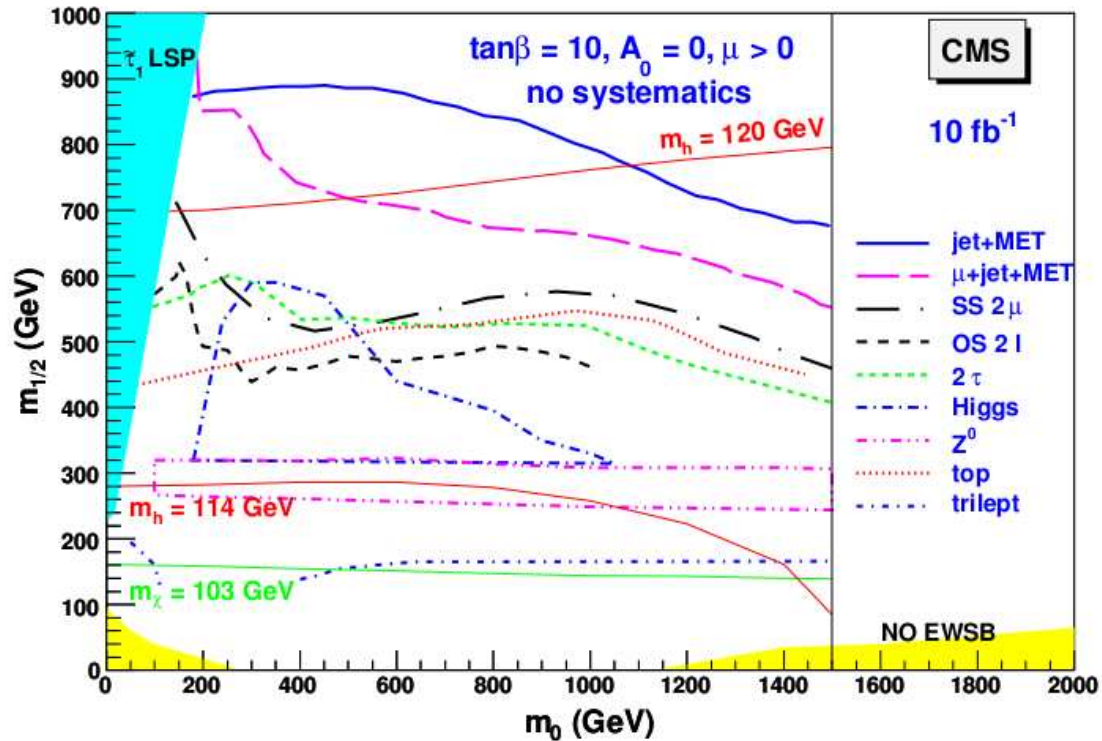


Amitava Datta

TIFR, 23/10/2009

# Introduction

## CMS Discovery Channels(mSUGRA)



Chargino mass reach :  $m_{\tilde{\chi}_1^\pm} \lesssim 140 \text{ GeV}$  for  $m_{\tilde{l}_R} \approx 100 \text{ GeV}$   
 $m_{\tilde{\chi}_1^\pm} \lesssim 110 \text{ GeV}$  for  $m_{\tilde{l}_R} > 500 \text{ GeV}$

# Introduction

---

## Points to be noted :

- The chargino mass reach is not much larger than the current LEP lower bounds.
- Almost the entire parameter space accessible to the clean trilepton signal is forbidden by the lower limit on  $m_h > 114.7$  GeV.
- In addition recent studies (Zack Sullivan, Edmond L. Berger, PRD 78,034030,(2008)) indicates there are backgrounds neglected in previous analyses. For e.g.  $Zb\bar{b}$ ,  $W\gamma^*/Z^*$ .

# PLAN OF THE TALK

---

- Can  $2l+1\tau$  and  $1l+2\tau$  produce visible signals  
if the  $3l$  signal is too weak
- Can  $3l + (2l+1\tau) + (1l+2\tau)$  signal can improve the discovery reach
- The status of the  $\tau$  rich final states in the mSUGRA model with low  $m_0 - m_{1/2}$

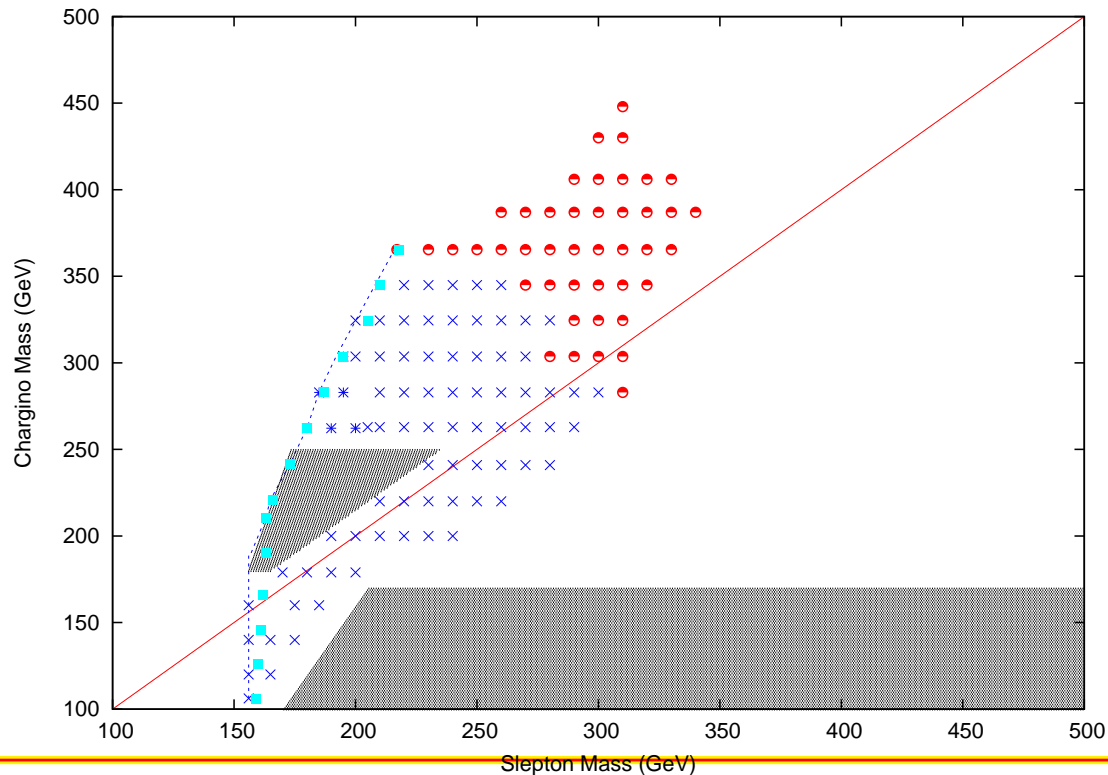


# Large mixing case

## Scenario MSSM: Heavy Squark, gluino

$M_1 \approx 0.5M_2$ ,  $M_{\tilde{g},\tilde{q}} = 3000\text{GeV}$ ,  $m_{\tilde{l}_L} = m_{\tilde{l}_R} = m_{\tilde{\ell}}$ ,  $A_0 = 0$ ,  
 $\tan\beta = 10$ ,  $m_A = 1000\text{GeV}$  and  $\mu = 1000$ . All values given at  
EWSB scale.

hep-ph/0906.1460

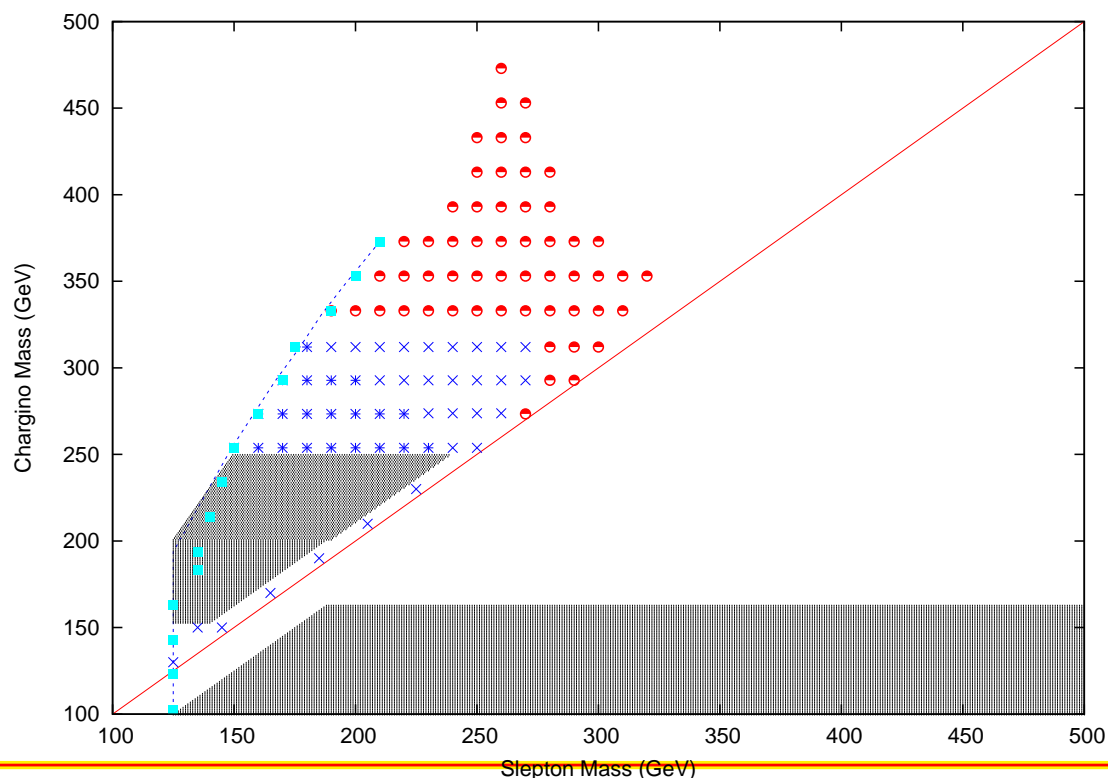


# Small mixing case

## Scenario MSSM: Heavy Squark, gluino

$M_1 \approx 0.5M_2$ ,  $M_{\tilde{g},\tilde{q}} = 3000\text{GeV}$ ,  $m_{\tilde{l}_L} = m_{\tilde{l}_R} = m_{\tilde{\ell}}$ ,  $A_0 = 0$ ,  
 $\tan\beta = 10$ ,  $m_A = 1000\text{GeV}$  and  $\mu = 500$ . All values given at  
EWSB scale.

hep-ph/0906.1460



# A representative point

**Ex.**  $M_1 = 125\text{GeV}$ ,  $M_2 = 250\text{GeV}$ ,  $M_{\tilde{g}} = 3000\text{GeV}$ ,  
 $m_{\tilde{\ell}} = 250\text{GeV}$ ,  $m_{\tilde{q}} = 3000\text{GeV}$ ,  $A_0 = 0$ ,  $\tan\beta = 10$ ,  
 $\mu = 500\text{GeV}$ ,  $m_A = 1000\text{GeV}$ . All values given at EWSB scale.

$\tilde{\chi}_1^0$	<b>123.6</b>	$\tilde{\tau}_1$	<b>236.1</b>	$\tilde{\nu}_{l_L}$	<b>241.8</b>
$\tilde{\nu}_{\tau_L}$	<b>241.8</b>	$\tilde{l}_R$	<b>253.8</b>	$\tilde{l}_l$	<b>254.1</b>
$\tilde{\chi}_1^\pm$	<b>254.8</b>	$\tilde{\chi}_2^0$	<b>254.9</b>	$\tilde{\tau}_2$	<b>270.6</b>
$\tilde{\chi}_3^0$	<b>518.9</b>	$\tilde{\chi}_4^0$	<b>531.7</b>	$\tilde{\chi}_2^\pm$	<b>531.9</b>
$\tilde{b}_1$	<b>2888.7</b>	$\tilde{q}_L$	<b>2890.0</b>	$\tilde{q}_R$	<b>2890.0</b>
$\tilde{b}_2$	<b>2893.0</b>	$\tilde{t}_1$	<b>2923.3</b>	$\tilde{g}$	<b>2942.9</b>
$\tilde{t}_2$	<b>2956.5</b>				

3/ Signal Non-Universal Gaugino Masses (S.Bhattacharya,Asesh K Datta and B. Mukhopadhyay; arXiv:0809.02012[hep-ph])

# BRs of the representative point

---

Decay modes	BR
$\tilde{\chi}_1^+ \rightarrow \tilde{\nu}_l l^+$	42.0
$\tilde{\chi}_1^+ \rightarrow \tilde{\nu}_\tau \tau^+$	21.4
$\tilde{\chi}_1^+ \rightarrow \tilde{l}_l \nu_l$	–
$\tilde{\chi}_1^+ \rightarrow \tilde{\tau}_1^+ \nu_\tau$	22
$\tilde{\chi}_2^0 \rightarrow \tilde{l}_l l$	–
$\tilde{\chi}_2^0 \rightarrow \tilde{\tau}_1 \tau$	25
$\tilde{\chi}_2^0 \rightarrow \tilde{\nu} \nu$	62

# Is this scenario dark matter allowed

$$\Omega_{CDM}h^2 = .0903$$

Channels	Relative Contribution
$\tilde{\chi}_1^0 \tilde{\chi}_1^0 \rightarrow \tau\tau$	16
$\tilde{\chi}_1^0 \tilde{\chi}_1^0 \rightarrow ll$	18
<b>Bulk</b>	<b>34</b>
$\tilde{\chi}_1^0 \tilde{\tau}_1 \rightarrow \tau h$	2
$\tilde{\chi}_1^0 \tilde{\tau}_1 \rightarrow Z\tau$	8
$\tilde{\chi}_1^0 \tilde{\tau}_1 \rightarrow A\tau$	22
$\tilde{\chi}_1^0 \tilde{\tau}_1 \rightarrow W\nu_\tau$	8
<b>Co-annihilation</b>	<b>40</b>

# Details of analysis

---

- $\tau - jets$  :  $E_{visible} = E_T^\tau - E_T^{\nu_\tau}$  ;  $|\eta_\tau| < 2.4$ . Tagging efficiency ( $\epsilon_\tau$ ) from  $E_{visible} = 30\text{GeV}$  given in CMS TDR-II.
- Leptons ( $l = e, \mu$ ):  $P_T^e \geq 17\text{ GeV}$ ,  $P_T^\mu \geq 10\text{ GeV}$  and  $|\eta^l| < 2.4$ . Lepton-jet isolation:  $\Delta R(l, j) > 0.5$ .
- $2\tau + 1l$  :  $\cancel{E}_T > 100\text{ GeV}$
- $1\tau + 2l$  :  $\cancel{E}_T > 100\text{ GeV}$  and  $80 < M_{inv}^{ll} < 100\text{ GeV}$ .

# Background (hep-ph/0906.1460)

---

	Signal	QCD	WZ	ZZ	$t\bar{t}$	Z + j
$\sigma$ (pb)	0.896	$1.48 \times E^8$	31.47	13.11	400	$1.7 \times E^5$
$2 \tau + 1 l$	0.00134	105.7	0.0066	0.0017	0.004	4.7
$\cancel{E}_T > 100$ GeV	0.00108	–	–	–	–	–
$1 \tau + 2 l$	0.00206	63.4	0.015	0.0111	0.24	19.6
$\cancel{E}_T > 100$ GeV	0.00072	–	0.0094	0.0026	.068	–
$80 < M_{inv}^{ll} < 100$ GeV	0.00045	–	–	–	.064	–

$\sigma \times \epsilon$  is given above.

The  $W\gamma^*/Z^*$  is numerically significant (0.0005 for  $2\tau + 1l$  and 0.0003 for  $1\tau + 2l$ ) but not negligible.

# GUT representative point

$M_1 = M_2 = 300, M_3 = 1200, m_{\tilde{\ell}} = 150, m_0 = 160, A_0 = 0,$   
 $\tan\beta = 10$  and  $\Omega(\Omega) = 0.10.$

$\tilde{\chi}_1^0$	<b>117.2</b>	$\tilde{\tau}_1$	<b>132.1</b>	$\tilde{l}_R$	<b>193.6</b>
$\tilde{\nu}_{\tau L}$	<b>221.6</b>	$\tilde{\nu}_{l L}$	<b>222.5</b>	$\tilde{\chi}_1^\pm$	<b>225.3</b>
$\tilde{\chi}_2^0$	<b>225.3</b>	$\tilde{l}_l$	<b>235.7</b>	$\tilde{\tau}_2$	<b>272.7</b>
$\tilde{\chi}_3^0$	<b>1433.5</b>	$\tilde{\chi}_4^0$	<b>1435.4</b>	$\tilde{\chi}_2^\pm$	<b>1436.1</b>
$\tilde{t}_1$	<b>1913.1</b>	$\tilde{b}_1$	<b>2101.9</b>	$\tilde{t}_2$	<b>2125.9</b>
$\tilde{b}_2$	<b>2240.9</b>	$\tilde{q}_L$	<b>2245.1</b>	$\tilde{q}_R$	<b>2248.6</b>
$\tilde{g}$	<b>2594.7</b>				



- 
- Intuitively this scenario can also arise in GMSB model.
  - In GMSB model mass arises via gauge interaction; hence strongly interacting sparticles could be expected to be much heavier than the sparticles in EW sector.

# Signals in mSUGRA

---

- Low  $m_0 - m_{1/2}$  is disfavoured in mSUGRA if  $A_0 = 0$  by Higgs mass bound.
- For  $A_0 \neq 0$ , there are regions of parameter space satisfying the Higgs mass bound as well as DM data.
- The three **benchmark scenarios** in mSUGRA are as follows :
  - **A:**  $m_0 = 120, m_{1/2} = 300, A_0 = -930, \tan\beta = 10, \mu > 0$
  - **B:**  $m_0 = 120, m_{1/2} = 350, A_0 = -930, \tan\beta = 10, \mu > 0$
  - **C:**  $m_0 = 120, m_{1/2} = 500, A_0 = 0, \tan\beta = 10, \mu > 0$

# Signals in mSUGRA

---

The dominant DM relic density producing mechanisms in the above scenarios are as follows:

- Scenario A is characterized by both **LSP pair annihilation** and **LSP- $\tilde{\tau}_1$  coannihilation**.
- In scenario B **LSP- $\tilde{\tau}_1$  coannihilation** dominates.
- In scenario C **LSP- $\tilde{\tau}_1$  coannihilation** is the only DM producing mechanism.

# Signals in mSUGRA

---

	A	B	C
$\sigma$ (pb)	0.747	0.403	0.106
$2 \tau + 1 l$	0.000276	0.000189	0.000210
$\cancel{E}_T > 100 \text{ GeV}$	0.000149	0.00010	0.00012
$1 \tau + 2 l$	0.000366	0.000387	0.001257
$\cancel{E}_T > 100 \text{ GeV}$	0.000187	0.000185	0.000831
$80 < M_{inv}^{ll} < 100 \text{ GeV}$	0.000172	0.000165	0.000670

None of them leads to observable signal at  $10 fb^{-1}$

# Signals in mSUGRA

---

Reasons for small (non observable) signals in mSUGRA :

- General model yeild a larger  $\tilde{\chi}_1^\pm - \tilde{\chi}_2^0$  production cross-section compared to mSUGRA for the same  $m_{\tilde{\chi}_1^\pm}$ .
- In mSUGRA with a common scalar mass  $m_0$  at  $M_G$ ,  $\tilde{\tau}_R$  becomes lightest charged slepton at weak scale.
- Decay products of  $\tilde{\tau}_1$  are soft.

# CONCLUSION

---

**Final states with tagged  $\tau - jets$  have excellent potential in the context of SUSY searches at the LHC.**