## Baryon number violating processes

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[^0]Effective operators: quark level....assuming perturbative new physics But is it completely understood ? Will come back to it at the end


Julian Heeck and Volodymyr
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Quark level dim-6

$$
\mathcal{O}_{d}^{(1)}=\left(\bar{d}_{i R}^{c} u_{j R}\right)\left(\bar{u}_{k L}^{c} e_{d L}-\bar{d}_{k L}^{c} \nu_{d L}\right) \epsilon_{i j k},
$$

$$
\mathcal{O}_{d}^{(2)}=\left(\bar{d}_{i L}^{c} u_{j L}\right)\left(\bar{u}_{k R}^{c} e_{d R}\right) \epsilon_{i j k},
$$

$$
\mathcal{O}_{d}^{(3)}=\left(\bar{d}_{i L}^{c} u_{j L}\right)\left(\bar{u}_{k L}^{c} e_{d L}-\bar{d}_{k L}^{c} \nu_{d L}\right) \epsilon_{i j k},
$$

$$
\mathcal{O}_{d}^{(4)}=\left(\bar{d}_{i R}^{c} u_{j R}\right)\left(\bar{u}_{k R}^{c} e_{d R}\right) \epsilon_{i j k} .
$$

## Hadron level

$$
\begin{aligned}
& \mathcal{O}_{d}^{(1)}=\alpha\left(\bar{e}_{d L}^{c} \operatorname{Tr} \mathcal{F} \xi B_{L} \xi-\bar{\nu}_{d L}^{c} \operatorname{Tr} \mathcal{F}^{\prime} \xi B_{L} \xi\right) \\
& \mathcal{O}_{d}^{(2)}=\alpha \bar{e}_{d R}^{c} \operatorname{Tr} \mathcal{F} \xi^{\dagger} B_{R} \xi^{\dagger} \\
& \mathcal{O}_{d}^{(3)}=\beta\left(\bar{e}_{d L}^{c} \operatorname{Tr} \mathcal{F} \xi B_{L} \xi^{\dagger}-\bar{\nu}_{d L}^{c} \operatorname{Tr} \mathcal{F}^{\prime} \xi B_{L} \xi^{\dagger}\right), \\
& \mathcal{O}_{d}^{(4)}=\beta \bar{e}_{d R}^{c} \operatorname{Tr} \mathcal{F} \xi^{\dagger} B_{R} \xi
\end{aligned}
$$

$$
\langle 0| \epsilon_{i j k}\left(u^{i T} C P_{R} d^{j}\right) P_{L} u^{k}\left|p^{(s)}\right\rangle=\alpha P_{L} u^{(s)} .
$$

$$
\langle 0| \epsilon_{i j k}\left(u^{i T} C P_{L} d^{j}\right) P_{L} u^{k}\left|p^{(s)}\right\rangle=\beta P_{L} u^{(s)} .
$$

$$
\begin{aligned}
\mathcal{L}_{d=6}= & y_{a b c d}^{1} \epsilon^{\alpha \beta \gamma}\left(\bar{d}_{a, \alpha}^{C} u_{b, \beta}\right)\left(\bar{Q}_{i, c, \gamma}^{C} \epsilon_{i j} L_{j, d}\right) \\
& +y_{a b c d^{\alpha} \epsilon^{\alpha \beta \gamma}}\left(\bar{Q}_{i, a, \alpha}^{C} \epsilon_{i j} Q_{j, b, \beta}\right)\left(\bar{u}_{c, \gamma}^{C} \ell_{d}\right) \\
& +y_{a b c c}^{3} \epsilon^{\alpha \beta \gamma} \epsilon_{i l} \epsilon_{j k}\left(\bar{Q}_{i, a, \alpha}^{C} Q_{j, b, \beta}\right)\left(\bar{Q}_{k, c, \gamma}^{C} L_{l, d}\right) \\
& +y_{a b c c d}^{4} \epsilon^{\alpha \beta \gamma}\left(\bar{d}_{a, \alpha}^{C} u_{b, \beta}\right)\left(\bar{u}_{c, \gamma}^{C} \ell_{d}\right)+\text { h.c. },
\end{aligned}
$$

$$
\Gamma\left(p \rightarrow e^{+} \pi^{0}\right) \simeq \frac{1}{2 \times 10^{34} \mathrm{yr}}\left|\frac{y_{1111}^{j}}{\left(3 \times 10^{15} \mathrm{GeV}\right)^{-2}}\right|^{2}
$$

## Processes

- $\underset{\text { (Benjamín Grinstein) }}{p} \mu^{+} / e^{+}, p \rightarrow K^{+} \bar{\nu}$ Lepto-Quark and diquark
- $p p \rightarrow \mu^{+} \mu^{+} / e^{+} e^{+}{ }_{\text {Lepto-quark and diquarks (K. S. Babu, R. }}$ N. Mohapatra, et al ), Weyl-Majorana( in preparation, MTA)


## INO ? 50 kiloton magnetized ICAL detector at the INO for 10 years

Ferjus Experiment with 900 ton iron tracking calorimeter?

- $n-\bar{n}$ six-dimensions (DC, MTA), Linear-moose model (MTA), Leptoquark and diquarks (K. S. Babu, R. N. Mohapatra, C. Grojean, J. D. Wells )
- Exotic Induced Nucleon Decays (IND)
- $D M+p \rightarrow n+e^{+} / \mu^{+}$Dark Majorana fermion (Amarjit Soni)
- $D M+p \rightarrow D M+e^{+} / \mu^{+}$Baryonic DM(Hooman Davoudiasl, Nikita Blinov), Weyl-Majorana interaction manuscript (in preparation MTA)
- $D M+p \rightarrow D M+\bar{p}+\mu^{+} \mu^{+} / e^{+} e^{+}$ Weyl-Majorana interaction manuscript (in preparation MTA)
- $\nu n \rightarrow \bar{\nu} \bar{n}$


## Proposed liquid scintillator ? Similar to SNO

$$
\Gamma(A \longrightarrow(A-2)+\text { mesons })
$$

- $n \rightarrow n^{\prime}$ mirror neutron models (Z. Berezhiani, D. McKeen, M. Pospelov), Linear-moose model (MTA) (Purely non-perturbative confining)

- $e^{-}+N(A, Z) \rightarrow e^{+}+N^{\prime}(A, Z-2)$

Inclusive Nucleon Decay Searches as a Frontier of Baryon Number Violation Julian Heeck ${ }^{1, *}$ and Volodymyr Takhistov ${ }^{2, \dagger}$

## Hylogenesis:

A Unified Origin for Baryonic Visible Matter and Antibaryonic Dark Matter
Hooman Davoudiasl, ${ }^{1}$ David E. Morrissey, ${ }^{2}$ Kris Sigurdson, ${ }^{3}$ and Sean Tulin ${ }^{2}$

## Baryon Destruction by Asymmetric Dark Matter

Hooman Davoudiasl ${ }^{(a)}$, David E. Morrissey ${ }^{(b)}$, Kris Sigurdson ${ }^{(c)}$, Sean Tulin ${ }^{(b)}$

Neutron oscillation and Baryogenesis from six dimensions
Mathew Thomas Arun ${ }^{1}$, Debajyoti Choudhury ${ }^{2}$

Baryon number violation from confining New Physics
Mathew Thomas Arun

Distinguishing $4 k+2$ dimensions with assisted baryon number violation

Dark Matter Antibaryons from a Supersymmetric Hidden Sector

Chiral. Lagrangiain for deep mine physics*

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        Mark Claudson and Mark B. Wise }\mp@subsup{}{}{\dagger
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                                    ABSTRACT
                                    The chiral Lagranitian for baryon number
    violating nucleon decay is derived and applied
to nucleon decays into strange and nonstrange
final states. The uncertainties in our pre-
dictions are discussed.

## Hidden MeV-scale dark matter in neutrino detectors

Jennifer Kile* and Amarjit Soni ${ }^{\dagger}$

Baryon number and lepton universality violation in leptoquark and diquark models

Nima Assad, Bartosz Fornal *, Benjamín Grinstein
decays:
(Induced) Proton decay

- $p \rightarrow e^{+}+\pi^{0}$
- $D M+p \rightarrow D M+e^{+}$
- $D M+p \rightarrow D M+e^{+}+\pi^{0}$
- Baryonic Dark Matter

|  | $S U(3)_{C}$ | $S U(2)_{L}$ | $U(1)_{Y}$ | $U(1)_{L} U(1)_{B}$ |  |
| :---: | :---: | :---: | :---: | :---: | :---: |
| $X$ | $\mathbf{1}$ | $\mathbf{1 ( 2 )}$ | $1 / 2$ | 0 | 1 |
| $\phi$ | $\mathbf{1}$ | $\mathbf{1}$ | 0 | $1 / 2$ | $1 / 2$ |
| $\Phi_{e}$ | $\mathbf{1}$ | $\mathbf{1}$ | 0 | 1 | 1 |

$$
\Gamma\left(p \rightarrow e^{+} \pi^{0}\right) \simeq \frac{1}{2 \times 10^{34} \mathrm{yr}}\left|\frac{y_{1111}^{j}}{\left(3 \times 10^{15} \mathrm{GeV}\right)^{-2}}\right|^{2} \quad \tau_{\mathrm{eff}}=1.5 \times 10^{33} \mathrm{yr}\left(\frac{0.7 \times 10^{-40} \mathrm{~cm}^{3} / \mathrm{s}}{(\sigma v)_{\mathrm{IND}}}\right)
$$

decays:
(Induced) Proton Proton annihilation / Hydrogen-antiHydrogen oscillation

- $p+p \rightarrow \mu^{+}+\mu^{+} / e^{+}+e^{+}$ (dim-12)
- $D M+p \rightarrow D M+\bar{p}+\mu^{+} \mu^{+} / e^{+} e^{+}$
- There could be striking signatures
- They arise from different operators
- How do you distinguish
- This is done by looking at the $\mu^{+} \mu^{+} / e^{+} e^{+}$in the final state.
- $\Lambda \gtrsim 1.5 \mathrm{Te} V$ for perturbative new physics

$V^{(1,0)}$
$\Gamma_{16}^{\mathcal{O}_{4}}{ }^{14}{ }^{14} C \ell^{+} \ell^{+}+\frac{21}{4 \pi^{2}}\left(\frac{m_{p}^{2}}{\Lambda_{4}^{4}}\right) r^{-3}$
$=\left(6.4 \times 10^{27} \mathrm{yrs}^{-1}\right)\left(\frac{\mathrm{GeV}}{\Lambda_{4}}\right)^{4}$


## Exotic decay: <br> $\mu^{-}+N(A, Z) \rightarrow e^{+}+N^{\prime}(A, Z-2)$


M. Lee and M.~MacKenzie,

Universe 8, no.4, 227 (2022) arXiv:2110.07093

$$
n-\bar{n} \quad\langle\bar{n}| \mathcal{H}_{\Delta B=2}|n\rangle=-\frac{1}{2} \epsilon \nu_{\bar{n}}^{T} C u_{n}
$$

## Confining New

## Physics

## Six-dimensions

$$
d s_{6}^{2}=b^{2}\left(x_{5}\right)\left[a^{2}\left(x_{4}\right) \eta_{\mu \nu} d x^{\mu} d x^{\nu}+d x_{4}^{2}\right]+d x_{5}^{2}
$$

$$
\mathcal{L}=\sqrt{-g_{6}}\left(M_{6}^{4} R_{6}-\Lambda_{6}\right)
$$

$$
+\sqrt{-g_{5}}\left[V_{1}\left(x_{5}\right) \delta\left(x_{4}\right)+V_{2}\left(x_{5}\right) \delta\left(x_{4}-\pi R_{y}\right)\right]
$$

$$
+\sqrt{-\tilde{g}_{5}}\left[V_{3}\left(x_{4}\right) \delta\left(x_{5}\right)+V_{4}\left(x_{4}\right) \delta\left(x_{5}-\pi r_{z}\right)\right]
$$

$$
\mathcal{L}_{i n t}=\eta_{u d} \phi \overline{u^{c}} d+\eta_{u e} \phi^{*} \overline{u^{c}} e+\zeta_{d d} \omega^{*} \overline{d^{c}} d
$$

where

$$
\begin{align*}
\zeta_{d d} & =\left(2 r_{z}\right)^{-1} z_{d d} \int d x_{5} b^{2} \chi_{\phi} \chi_{d}^{2} \\
\eta_{u d} & =\left(2 r_{z}\right)^{-1} y_{u d} \int d x_{5} b^{2} \chi_{\phi} \chi_{u} \chi_{d} \\
\rho & =\left(2 r_{z}\right)^{-1} \lambda \int d x_{5} b^{2} \chi_{\phi}^{2} \chi_{\omega}  \tag{6}\\
\eta_{u e} & =y_{u e} b^{2}(0) \chi_{\phi}(0) \chi_{u}(0)
\end{align*}
$$



$$
+\rho M \phi^{2} \omega+\text { h.c. },
$$

$$
\begin{aligned}
\mathcal{L} & =i \bar{n} \gamma^{\mu} \partial_{\mu} n-\frac{m_{n}}{2}\left[\bar{n} n+\bar{n}^{c} n^{c}\right]-\frac{\epsilon}{2}\left[\bar{n}^{c} n+\bar{n} \overline{n^{c}}\right] \\
\mathcal{V}_{n} & =\sum_{i=1}^{N}\left(M n_{(i)}^{\prime c} n_{(i-1)}^{\prime}-m n_{(i)}^{\prime c} n_{(i)}^{\prime}\right)-y v \bar{n} n_{(0)}^{\prime} \\
& +\mathcal{L}_{i n t},
\end{aligned}
$$

| $\kappa$ | 1 | 0.1 | 0.01 | 0.001 |
| :---: | :---: | :---: | :---: | :---: |
| $m_{\phi}(\mathrm{TeV})=m_{\omega}=M$ | 670 | 106 | 17 | 2.6 |
| $m_{\phi}(\mathrm{TeV})=m_{\omega} / 3=M / 3$ | 345 | 55 | 9 | 1.4 |

$$
\begin{align*}
\mathcal{L} & =i \bar{n} \gamma^{\mu} \partial_{\mu} n-\frac{m_{n}}{2}\left[\bar{n} n+\bar{n}^{c} n^{c}\right]-\frac{\left(y_{\text {eff }} v\right)^{2}}{m_{M}} \bar{n}^{c} n \\
& +\mathcal{L}_{\text {int }}+\text { h.c. } \tag{10}
\end{align*}
$$

# (induced) $p p \rightarrow \mu^{+} \mu^{+}$: <br> At INO 50 Kiloton ( $\sim 10^{\wedge} 9$ Avogadro number) $? e^{+} e^{+}$(with scintillators) Frejus experiment? 

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Physics Letters B 269 (1991) 227-233 North-Holland
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PHYSICS LETTERS B
$\qquad$

Lifetime limits on ( $B-L$ )-violating nucleon decay and di-nucleon decay modes from the Fréjus experiment

Fréjus Collaboration

## 2. The experiment

The Fréjus detector has been described in detail in ref. [19]. The fine granularity of this 900 ton tracking calorimeter is achieved by a sandwich structure consisting of 912 flash chambers ( $5 \mathrm{~mm} \times 5 \mathrm{~mm}$ cells) and iron ( 3 mm ) planes interspersed with 113 planes of Geiger tubes ( $15 \mathrm{~mm} \times 15 \mathrm{~mm}$ cells) which pro vide the trigger. The detector cells are oriented vertical and horizontal alternately, thus providing two in dependent orthogonal views for each event.
The trigger requires grouped hits in a small volume ( $1 \mathrm{~m}^{3}$ ) typical for nucleon decay events corresponding to an energy threshold of about 200 MeV . For the
baryon-number-violating processes investigated here this trigger threshold results in a trigger efficiency ranging from $25 \%$ for $\mathrm{p} \rightarrow \mu^{+} v v$ to $97 \%$ for $\mathrm{pp} \rightarrow \mathrm{e}^{+} \mathrm{e}^{+}$ The efficiency of the detector was constantly moniored by analyzing the atmospheric muons passing hrough the apparatus. The average trigger rate is 45 vents per hour. Half of the triggers are due to cosmic ray muons while the rest is induced by local radioac tivity and electronic noise. The event rate produced by interactions of atmospheric neutrinos is about one event per week.

## UNDERGROUND Frejus tunnel experiment

To search for proton instability and to measure the proton lifetime, an Orsay / Saclay / Ecole Polytechnique collaboration last year put forward a proposal for an experiment in the Frejus tunnel using a calorimetric detector with a fiducial mass of 1000 tons. A German group from Wuppertal added its support to the proposed experiment, which was approved in May.
The detector, which has a total mass of 1500 tons, measures $6 \times 6 \times$ 20 m . The nucleon source consists of iron plates 3 or 4 mm thick. The fine grain is provided by 1600 banks, measuring $6 \times 6 \mathrm{~m}$, of polypropylene plasma tubes 6 mm long, similar to those operating in a Fermilab neu-

| $\Delta(B-L)$ | $\Delta B=2$ | (\%) | B | $N_{\text {c }}$ | $S_{90}$ | $\begin{aligned} & \tau_{\mathrm{N}}^{\prime} / \mathrm{BR} \\ & \left(10^{30} \mathrm{yr}\right) \end{aligned}$ | $\begin{aligned} & \tau_{\mathrm{N}} / \mathrm{BR} \\ & \left(10^{30} \mathrm{yr}\right) \end{aligned}$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 2 | $\mathrm{pp} \rightarrow \pi^{+} \pi^{+}$ | 19.0 | 2.34 | 4 | 5.81 | 0.5 | 0.7 |
|  | $\mathrm{pn} \rightarrow \pi^{+} \pi^{0}$ | 23.4 | 0.31 | 0 | 2.30 | 2.0 | 2.0 |
|  | $\mathrm{nn} \rightarrow \pi^{0} \pi^{0}$ | 36.7 | 0.78 | 0 | 2.30 | 3.4 | 3.4 |
|  | $\mathrm{nn} \rightarrow \pi^{+} \pi^{-}$ | 19.8 | 2.18 | 4 | 5.94 | 0.5 | . 0.7 |
| 0 | $\mathrm{pp} \rightarrow \mathrm{e}^{+} \mathrm{e}^{+}$ | 62.3 | <0.10 | 0 | 2.30 | 5.8 | 5.8 |
|  | $\mathrm{pp} \rightarrow \mathrm{e}^{+} \mu^{+}$ | 38.8 | <0.10 | 0 | 2.30 | 3.6 | 3.6 |
|  | pp $\rightarrow \mu^{+} \mu^{+}$ | 18.4 | 0.62 | 0 | 2.30 | 1.7 | 1.7 |
|  | $\mathrm{pn} \rightarrow \mathrm{e}^{+} \overline{\mathrm{v}}$ | 52.8 | 9.67 | 5 | 3.77 | 1.1 | 2.8 |
|  | $\mathrm{pn} \rightarrow \mathbf{u}^{+} \overline{\mathrm{v}}$ | 37.4 | 4.37 | 4 | 4.61 | 0.9 | 1.6 |

trino experiment. These tubes are triggered by 185 Geiger tube banks of the same dimensions. With these it will be possible to measure the time of flight of the particles and to detect any particularly slow particles, such as monopoles. All these detector components are fitted in alternate array, giving the whole a modular structure. The iron sampling precision and the fineness of the grain will make it possible to discover the charge of the tracks and to measure their energy with a high degree of accuracy.
The experiment will be carried out in the Modane underground laboraory fitted out beside the Frejus tun nel linking Modane in France to Bar donecchia in Italy and located on French territory some 100 m from he border An initial excavation $800 \mathrm{~m}^{3}$ was made before the tunnel was completed in 1980 and back ground measurements were made.

They showed that the muon rate was indeed reduced by a factor of the order of $10^{6}$, as expected by the 1500 m of rock covering the chamber. This excavation will be enlarged at the beginning of 1982 to take a $3000 \mathrm{~m}^{3}$ laboratory. A start will be made on installing the detector at the end of 1982 Since it is modular it can begin operating before it is completed. It will be sensitive enough to allow the measurement of a lifetime of $10^{31}$ years in one year and to reach a lower limit of $10^{32}$ years in a few years.

## Dinucleon and Nucleon Decay to Two-Body Final States with no Hadrons in Super-Kamiokande

arXiv:1811.12430

The Super-Kamiokande (SK) water Cherenkov detector, with a fiducial volume of 22.5 kilotons, contains $1.2 \times 10^{34}$ nucleons. SK lies one kilometer under Mt. Ikenoyama in Japan's Kamioka Observatory. The detector is cylindrical with a diameter of 39.3 meters and a height of 41.4 meters, optically separated into an inner and an outer region. Eight-inch photomultiplier tubes (PMTs) line the outer detector facing outwards and serve primarily as a veto for cosmic ray muons, and 20 -inch PMTs face inwards to measure Cherenkov light in the inner detector [12].


FIG. 1. (color online) An SK event display of a typical $p p \rightarrow$ $e^{+} \mu^{+}$event shown in $\theta-\phi$ view. The non-showering ring (from the $\mu^{+}$) is on the left and the showering ring (from the $e^{+}$) is on the right. The energy of each ring is approximately 900 MeV .


The following selection criteria are applied to signal MC, atmospheric $\nu \mathrm{MC}$, and data:
(A1) Events must be fully contained in the inner detector with the event vertex within the fiducial volume (two meters inward from the detector walls),
(A2) There must be two Cherenkov rings,
(A3) Both rings must be showering for the $p p \rightarrow e^{+} e^{+}$, $n n \rightarrow e^{+} e^{-}, n n \rightarrow \gamma \gamma$ and $p \rightarrow e^{+} \gamma$ modes; one ring must be showering and one ring must be nonshowering for the $p p \rightarrow e^{+} \mu^{+}, n n \rightarrow e^{+} \mu^{-}, n n \rightarrow$ $e^{-} \mu^{+}$and $p \rightarrow \mu^{+} \gamma$ modes; both rings must be non-showering for the $p p \rightarrow \mu^{+} \mu^{+}, n n \rightarrow \mu^{+} \mu^{-}$ modes (see note in [19]),
(A4) There must be zero Michel electrons for the $p p \rightarrow$ $e^{+} e^{+}, n n \rightarrow e^{+} e^{-}, n n \rightarrow \gamma \gamma$ and $p \rightarrow e^{+} \gamma$ modes there must be less than or equal to one Michel electron for the $p p \rightarrow e^{+} \mu^{+}, n n \rightarrow e^{+} \mu^{-}, n n \rightarrow e^{-} \mu^{+}$ and $p \rightarrow \mu^{+} \gamma$ modes; there is no Michel electron cut for the $p p \rightarrow \mu^{+} \mu^{+}, n n \rightarrow \mu^{+} \mu^{-}$modes (see note in [20]),
(A5) The reconstructed total mass, $M_{t o t}$, should be $1600 \leq M_{t o t} \leq 2050 \mathrm{MeV} / \mathrm{c}^{2}$ for the dinucleon decay modes; the reconstructed total mass should be $800 \leq M_{t o t} \leq 1050 \mathrm{MeV} / \mathrm{c}^{2}$ for the nucleon decay modes,
(A6) The reconstructed total momentum, $P_{\text {tot }}$, should be $0 \leq P_{\text {tot }} \leq 550 \mathrm{MeV} / \mathrm{c}$ for the dinucleon decay modes; for the nucleon decay modes, it should be $100 \leq P_{t o t} \leq 250 \mathrm{MeV} / \mathrm{c}$ for the event to be in the "High $P_{\text {tot }}$ " signal box and $0 \leq P_{\text {tot }} \leq 100 \mathrm{MeV} / \mathrm{c}$ for the event to be in the "Low $P_{\text {tot }}$ " signal box,
(A7) [SK-IV nucleon decay searches only] There must be zero tagged neutrons

TABLE IV. Parameters of past (KAM [114, 115]), running (SK [116, 117]), and future (HK-3TankLD and HK-1TankHD) water Cherenkov detectors. The KAM and SK have undergone several configuration changes and parameters for KAM-II and SK-IV are referred in the table. The single-photon detection efficiencies are products of the quantum efficiency at peak ( $\sim 400 \mathrm{~nm}$ ), photo-electron collection efficiency, and threshold efficiency. Most right column (HK-1TankHD) shows another design under study which consist of one tank instrumented with high density PMTs.

|  | KAM | SK | HK-3TankLD | HK-1TankHD |
| :--- | :---: | :---: | :---: | :---: |
| Depth | $1,000 \mathrm{~m}$ | $1,000 \mathrm{~m}$ | 650 m | 650 m |
| Dimensions of water tank |  |  |  |  |
| diameter | $15.6 \mathrm{~m} \phi$ | $39 \mathrm{~m} \phi$ | $74 \mathrm{~m} \phi$ | $74 \mathrm{~m} \phi$ |
| $\quad$ height | 16 m | 42 m | 60 m | 60 m |
| Total volume | 4.5 kton | 50 kton | 774 kton | 258 kton |
| Fiducial volume | 0.68 kton | 22.5 kton | 560 kton | 187 kton |
| Outer detector thickness | $\sim 1.5 \mathrm{~m}$ | $\sim 2 \mathrm{~m}$ | $1 \sim 2 \mathrm{~m}$ | $1 \sim 2 \mathrm{~m}$ |
| Number of PMTs |  |  |  |  |
| $\quad$ inner detector (ID) | $948(50 \mathrm{~cm} \phi)$ | $11,129(50 \mathrm{~cm} \phi)$ | $40,000(50 \mathrm{~cm} \phi)$ | $40,000(50 \mathrm{~cm} \phi)$ |
| $\quad$ outer detector (OD) | $123(50 \mathrm{~cm} \phi)$ | $1,885(20 \mathrm{~cm} \phi)$ | $20,000(20 \mathrm{~cm} \phi)$ | $6,700(20 \mathrm{~cm} \phi)$ |
| Photo-sensitive coverage | $20 \%$ | $40 \%$ | $13 \%$ | $40 \%$ |
| Single-photon detection | unknown | $12 \%$ | $24 \%$ | $24 \%$ |
| efficiency of ID PMT |  |  |  | 1 nsec |
| Single-photon timing | $\sim 4 \mathrm{nsec}$ | $2-3 \mathrm{nsec}$ |  | 1 nsec |
| resolution of ID PMT |  |  |  |  |

## Super-Kamiokande and Hyper-Kamiokande (~650 m)

|  | Lifetime limit |  |
| :--- | :---: | :---: |
| Decay mode | $\begin{array}{c}\text { per oxygen nucleus } \\ \left(\times 10^{33}\right.\end{array}$ years $)$ | per nucleon <br> $\left(\times 10^{34}\right.$ years $)$ |
| $p p \rightarrow e^{+} e^{+}$ | 4.2 | - |
| $n n \rightarrow e^{+} e^{-}$ | 4.2 | - |
| $n n \rightarrow \gamma \gamma$ | 4.1 | - |
| $p p \rightarrow e^{+} \mu^{+}$ | 4.4 | - |
| $n n \rightarrow e^{+} \mu^{-}$ | 4.4 | - |
| $n n \rightarrow e^{-} \mu^{+}$ | 4.4 | - |
| $p p \rightarrow \mu^{+} \mu^{+}$ | 4.4 | - |
| $n n \rightarrow \mu^{+} \mu^{-}$ | 4.4 | 4.1 |
| $p \rightarrow e^{+} \gamma$ | - | 2.1 |
| $p \rightarrow \mu^{+} \gamma$ | - |  |

arXiv:1811.12430 (Dinucleon and Nucleon Decay to
Two-Body Final States with no Hadrons in Super-
Kamiokande )

| Mode | Sensitivity $(90 \%$ CL $)$ [years] | Current limit [years] |
| :--- | :---: | :---: |
| $p \rightarrow e^{+} \pi^{0}$ | $1.2 \times 10^{35}$ | $1.4 \times 10^{34}$ |
| $p \rightarrow \bar{\nu} K^{+}$ | $2.8 \times 10^{34}$ | $0.7 \times 10^{34}$ |
| $p \rightarrow \mu^{+} \pi^{0}$ | $9.0 \times 10^{34}$ | $1.1 \times 10^{34}$ |
| $p \rightarrow e^{+} \eta^{0}$ | $5.0 \times 10^{34}$ | $0.42 \times 10^{34}$ |
| $p \rightarrow \mu^{+} \eta^{0}$ | $3.0 \times 10^{34}$ | $0.13 \times 10^{34}$ |
| $p \rightarrow e^{+} \rho^{0}$ | $1.0 \times 10^{34}$ | $0.07 \times 10^{34}$ |
| $p \rightarrow \mu^{+} \rho^{0}$ | $0.37 \times 10^{34}$ | $0.02 \times 10^{34}$ |
| $p \rightarrow e^{+} \omega^{0}$ | $0.84 \times 10^{34}$ | $0.03 \times 10^{34}$ |
| $p \rightarrow \mu^{+} \omega^{0}$ | $0.88 \times 10^{34}$ | $0.08 \times 10^{34}$ |
| $n \rightarrow e^{+} \pi^{-}$ | $3.8 \times 10^{34}$ | $0.20 \times 10^{34}$ |
| $n \rightarrow \mu^{+} \pi^{-}$ | $2.9 \times 10^{34}$ | $0.10 \times 10^{34}$ |

KEK Preprint 2016-21
(HYPER-KAMIOKANDE design report )

| Mode | Sensitivity $(90 \% \mathrm{CL})$ [years] | Current limit [years] |
| :--- | :---: | :---: |
| $p \rightarrow e^{+} \nu \nu$ | $10.2 \times 10^{32}$ | $1.7 \times 10^{32}$ |
| $p \rightarrow \mu^{+} \nu \nu$ | $10.7 \times 10^{32}$ | $2.2 \times 10^{32}$ |
| $p \rightarrow e+X$ | $31.1 \times 10^{32}$ | $7.9 \times 10^{32}$ |
| $p \rightarrow \mu^{+} X$ | $33.8 \times 10^{32}$ | $4.1 \times 10^{32}$ |
| $n \rightarrow \nu \gamma$ | $23.4 \times 10^{32}$ | $5.5 \times 10^{32}$ |
| $n p \rightarrow e^{+} \nu$ | $6.2 \times 10^{32}$ | $2.6 \times 10^{32}$ |
| $n p \rightarrow \mu^{+} \nu$ | $4.2 \times 10^{32}$ | $2.0 \times 10^{32}$ |
| $n p \rightarrow \tau^{+} \nu$ | $6.0 \times 10^{32}$ | $3.0 \times 10^{32}$ |

## Proposal: <br> Neutron lifetime/ Double proton decay

A Proposed Radiochemical Approach to the Nucleon Lifetime

$\tau_{p} \geqslant \frac{10^{3} \mathrm{~kg} \times 68 \% \times 18 / 58}{1.66 \cdot 10^{-27} \mathrm{~kg} / \text { nucleon }} / 60 / \mathrm{y}$

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$\geqslant 2 \cdot 10^{27} \mathrm{y}$.
W. Maenhatt ( ${ }^{* *}$ )

Department of Chemistry, University of Maryland - College Park, Md. 20742
(ricevuto il 13 Agosto 1975)

- If one of the neutrons in ${ }^{58} \mathrm{Ni}$ should decay then ${ }^{57} \mathrm{Ni}$ would beta decay to ${ }^{57} \mathrm{Co}$
- Look for the appearance of this activity in a large mass of well-shielded Ni.
- We can also look for the process ${ }^{39} K \rightarrow{ }^{37} \operatorname{Ar}$ (Potassium acetate powder)
- Searching for dim-6,12


## Proposal:

$3 n \rightarrow 3 \nu$ dim-18 at quark level/ dim-9 hadron level
$n n \rightarrow \bar{n} \bar{\nu}$ dim-15 at quark level/ dim-6 hadron level

- High energy New Physics is out of question !!
- Could differentiate non-perturbative low-energy New Physics
- ${ }^{59} \mathrm{Co} \rightarrow{ }^{56} \mathrm{Co} \rightarrow{ }^{56} \mathrm{Fe}$
- Through 3n annihilation (probably no one has done this yet)
- 56 Fe give out 846.8 KeV gamma rays. Need scintillators
- 6.4 KeV X ray of Fe would then allow the determination of the number of atoms of 59Co at source.


## Summary:

- $\Delta B=\Delta L=2$ processes could indicate perturbative New Physics at 1.5 TeV
- Searches for DM induced nucleon decays
- Proposal: 50 Kton INO and liquid scintillators could already be in place to look for baryon number violating currents
- Flavour violating $\mu^{-}+N(A, Z) \rightarrow e^{+}+N^{\prime}(A, Z-2)$ could be interesting
- Proposal: ${ }^{58} \mathrm{Ni}$ should decay then ${ }^{57} \mathrm{Ni}$ would beta decay to ${ }^{57} \mathrm{Co}$. ${ }^{39} \mathrm{~K} \rightarrow{ }^{37} \mathrm{Ar}$ (Potassium acetate powder) to look for dim -6,12 processes
- Proposal: ${ }^{59} \mathrm{Co} \rightarrow{ }^{56} \mathrm{Co} \rightarrow{ }^{56} \mathrm{Fe}$ with 846.8 KeV gamma rays could differentiate non-perturbative low-energy New Physics
- Thank you for listening


## Beckuc Sildes

$$
\begin{aligned}
\mathcal{L}_{d=6}= & y_{a b c d}^{1} \epsilon^{\alpha \beta \gamma}\left(\overline{\bar{a}}_{a, \alpha}^{C} u_{b, \beta}\right)\left(\bar{Q}_{i, c, \gamma}^{C} \epsilon_{i j} L_{j, d}\right) \\
& +y_{a b c d}^{2} \epsilon^{\alpha \beta \gamma}\left(\bar{Q}_{i, a, \alpha}^{C} \epsilon_{i j} Q_{j, b, \beta}\right)\left(\bar{u}_{c, \gamma}^{C} \ell_{d}\right) \\
& +y_{a b c d}^{3} \epsilon^{\alpha \beta \gamma} \epsilon_{i l l} \epsilon_{j k}\left(\bar{Q}_{i, a, \alpha}^{C} Q_{j, b, \beta}\right)\left(\bar{Q}_{k, c, \gamma}^{C} L_{l, d}\right) \\
& +y_{a b c c}^{4} \epsilon^{\alpha \beta \gamma}\left(\bar{d}_{a, \alpha}^{C} u_{b, \beta}\right)\left(\bar{u}_{c, \gamma}^{C} \ell_{d}\right)+\text { h.c. }
\end{aligned}
$$

$$
\Gamma\left(p \rightarrow e^{+} \pi^{0}\right) \simeq \frac{1}{2 \times 10^{34} \mathrm{yr}}\left|\frac{y_{1111}^{j}}{\left(3 \times 10^{15} \mathrm{GeV}\right)^{-2}}\right|^{2}
$$

$$
\Gamma\left(n \rightarrow \bar{\nu}_{\tau} \pi^{0}\right) \simeq \frac{1}{10^{33} \mathrm{yr}}\left|\frac{y_{3333}^{4}}{\left(5 \times 10^{8} \mathrm{GeV}\right)^{-2}}\right|^{2}
$$

$$
\mathscr{L}_{\text {eff }}=\frac{1}{\Lambda_{\mathrm{IND}}^{3}} \times\left\{\begin{array}{ll}
\epsilon_{\alpha \beta \gamma}\left(d_{R}^{\alpha} s_{R}^{\beta}\right)\left(u_{R}^{\gamma} \Psi_{R}\right) \Phi & \text { (case I) }  \tag{52}\\
\epsilon_{\alpha \beta \gamma}\left(s_{R}^{\alpha} u_{R}^{\beta}\right)\left(d_{R}^{\gamma} \Psi_{R}\right) \Phi & \text { (case II) } \\
\epsilon_{\alpha \beta \gamma}\left(u_{R}^{\alpha} d_{R}^{\beta}\right)\left(s_{R}^{\gamma} \Psi_{R}\right) \Phi & \text { (case III) }
\end{array}, \quad \frac{1}{\Lambda_{\mathrm{IND}}^{3}} \equiv \sum_{a=1,2} \frac{2 \bar{\zeta}_{a}^{*} Z_{31} V_{11}^{*} b_{P} \lambda_{a}^{\prime} \lambda^{\prime}}{m_{\widetilde{P}_{1}}^{2} m_{\widetilde{P}_{2}}^{2} m_{X_{a}}}\right.
$$

Here, we have neglected higher derivative terms, and $\Lambda_{\text {IND }}$ characterizes the IND mass scale. The different cases, corresponding to different baryon transfer interactions in Eq. (12), lead to different fermion contractions.

The effective IND rate for nucleon $N=p, n$ is

$$
\begin{equation*}
\Gamma(N \rightarrow K)=n_{\Psi}(\sigma v)_{\mathrm{IND}}^{N \Psi \rightarrow K \Phi^{\dagger}}+n_{\Phi}(\sigma v)_{\mathrm{IND}}^{N \Phi \rightarrow K \bar{\Psi}} \tag{53}
\end{equation*}
$$

where $n_{\Psi, \Phi}$ are the local DM number densities and $(\sigma v)_{\text {IND }}$ is the IND cross section. The IND lifetime can be expressed as

$$
\begin{equation*}
\tau(N \rightarrow K)=\frac{1}{\Gamma(N \rightarrow K)}=\frac{(1+r)\left(\Omega_{\mathrm{DM}} / \Omega_{\mathrm{b}}\right) m_{p}}{2 \rho_{\mathrm{DM}}\left[r(\sigma v)_{\mathrm{IND}}^{N \Psi \rightarrow K \Phi^{\dagger}}+(\sigma v)_{\mathrm{IND}}^{N \Phi \rightarrow K \bar{\Psi}}\right]} \tag{54}
\end{equation*}
$$

with local DM mass density $\rho_{\mathrm{DM}}=m_{\Psi} n_{\Psi}+m_{\Phi} n_{\Phi}$, and assuming the local ratio $r \equiv n_{\Psi} / n_{\Phi}$ is the same as over cosmological scales. The IND cross section is estimated as

$$
\begin{equation*}
(\sigma v)_{\mathrm{IND}} \approx \frac{m_{\mathrm{QCD}}^{4}}{16 \pi \Lambda_{\mathrm{IND}}^{6}} \approx 10^{-39} \mathrm{~cm}^{3} / \mathrm{s} \times\left(\frac{\Lambda_{\mathrm{IND}}}{1 \mathrm{TeV}}\right)^{-6} \tag{55}
\end{equation*}
$$

with QCD scale $m_{\mathrm{QCD}} \approx 1 \mathrm{GeV} \cdot{ }^{11}$ For $r \sim \mathcal{O}(1)$, the IND lifetime is

$$
\begin{equation*}
\tau(N \rightarrow K) \approx 10^{32} \mathrm{yrs} \times\left(\frac{(\sigma v)_{\mathrm{IND}}}{10^{-39} \mathrm{~cm} / \mathrm{s}}\right)^{-1}\left(\frac{\rho_{\mathrm{DM}}}{0.3 \mathrm{GeV} / \mathrm{cm}^{3}}\right)^{-1} \tag{56}
\end{equation*}
$$

## Backup slides



## Backup slides



## Backup slides

## Table 1

The number of source particles for the decay (e.g. protons, neutrons or nuclei) per kiloton for the different baryon-number-violating processes analyzed in this paper. Since for the nucleon decay modes with virtual meson exchange one of the two nucleons simply acts as spectator we take the number of protons or neutrons as for ( $B-L$ ) -violating nucleon decay, except for decay channels of the type $\mathrm{pn} \rightarrow \ell^{+} \mathrm{N}$ where either the neutron or the proton may decay. We thus take the total number of nucleons in this case. The decay rate of the true di-nucleon decay modes should be strongly affected by nuclear physics. Independent of physical details the lifetime should be proportional to the nuclear density which is nearly independent of the nuclear size. We therefore use the number of iron nuclei to calculate the lower lifetime limits for di-nucleon decay channels.

| Process | $N_{\mathrm{d}}$ <br> $\left(10^{32} \mathrm{kt}^{-1}\right)$ |
| :--- | :--- |
| $\mathrm{p} \rightarrow \mathrm{e}^{-} \mathrm{X}$ | 2.80 |
| $\mathrm{n} \rightarrow \mathrm{e}^{-} \mathrm{X}$ | 3.23 |
| $\mathrm{pp} \rightarrow \ell^{+} \mathrm{N}$ | 2.80 |
| $\mathrm{pn} \rightarrow \ell^{+} \mathrm{N}$ | 6.03 |
| $\mathrm{nn} \rightarrow \ell^{+} \mathrm{N}$ | 3.23 |
| $\mathrm{NN} \rightarrow \ell^{+} \ell$ | 0.11 |
| $\mathrm{NN} \rightarrow \pi \pi$ | 0.11 |

## 

$$
\begin{aligned}
d s_{6}^{2} & =b^{2}\left(x_{5}\right)\left[a^{2}\left(x_{4}\right) \eta_{\mu \nu} d x^{\mu} d x^{\nu}+d x_{4}^{2}\right]+d x_{5}^{2} \\
\mathcal{L} & =\sqrt{-g_{6}}\left(M_{6}^{4} R_{6}-\Lambda_{6}\right) \\
& +\sqrt{-g_{5}}\left[V_{1}\left(x_{5}\right) \delta\left(x_{4}\right)+V_{2}\left(x_{5}\right) \delta\left(x_{4}-\pi R_{y}\right)\right] \\
& +\sqrt{-\tilde{g}_{5}}\left[V_{3}\left(x_{4}\right) \delta\left(x_{5}\right)+V_{4}\left(x_{4}\right) \delta\left(x_{5}-\pi r_{z}\right)\right] .
\end{aligned}
$$

$$
\begin{aligned}
\mathcal{L}_{i n t} & =\eta_{u d} \phi \overline{u^{c}} d+\eta_{u e} \phi^{*} \overline{u^{c}} e+\zeta_{d d} \omega^{*} \overline{d^{c}} d \\
& +\rho M \phi \phi^{2} \omega+h . c .
\end{aligned}
$$

where

$$
\begin{aligned}
\zeta_{d d} & =\left(2 r_{z}\right)^{-1} z_{d d} \int d x_{5} b^{2} \chi_{\phi} \chi_{d}^{2} \\
\eta_{u d} & =\left(2 r_{z}\right)^{-1} y_{u d} \int d x_{5} b^{2} \chi_{\phi} \chi_{u} \chi_{d} \\
\rho & =\left(2 r_{z}\right)^{-1} \lambda \int d x_{5} b^{2} \chi_{\phi}^{2} \chi_{\omega} \\
\eta_{u e} & =y_{u e} b^{2}(0) \chi_{\phi}(0) \chi_{u}(0)
\end{aligned}
$$



$\epsilon_{B} \approx \frac{M^{2}}{4 \pi^{2} m_{S}^{2} \beta} \log \left(\frac{x_{\omega}-x_{\phi}+1+\beta}{x_{\omega}-x_{\phi}+1-\beta}\right) \mathcal{A}_{\rho}$
$\mathcal{A}_{\rho} \equiv \operatorname{Im}\left[\widetilde{\rho}_{12}^{*}\left(\rho^{\dagger} \widetilde{\rho} \rho\right)_{12}\right] / \sum_{i, j}\left|\widetilde{\rho}_{i j}\right|^{2}$

## 



$$
\begin{align*}
\mathcal{L} & =i \bar{n} \gamma^{\mu} \partial_{\mu} n-\frac{m_{n}}{2}\left[\bar{n} n+\bar{n}^{c} n^{c}\right]-\frac{\epsilon}{2}\left[\bar{n}^{c} n+\bar{n} \bar{n}^{c}\right] \\
\mathcal{V}_{n} & =\sum_{i=1}^{N}\left(M n_{(i)}^{\prime c} n_{(i-1)}^{\prime}-m n_{(i)}^{\prime c} n_{(i)}^{\prime}\right)-y v \bar{n} n_{(0)}^{\prime}  \tag{10}\\
& +\mathcal{L}_{i n t},
\end{align*}
$$

$$
\mathcal{L}=i \bar{n} \gamma^{\mu} \partial_{\mu} n-\frac{m_{n}}{2}\left[\bar{n} n+\bar{n}^{c} n^{c}\right]-\frac{\left(y_{e f f} v\right)^{2}}{m_{M}} \overline{n^{c}} n
$$

$$
+\mathcal{L}_{i n t}+\text { h.c. } .
$$

$$
\langle\bar{n}| \mathcal{H}_{\Delta B=2}|n\rangle=-\frac{1}{2} \epsilon \nu_{\bar{n}}^{T} C u_{n}
$$

$$
\begin{aligned}
\sigma_{n^{0} \mathcal{N}} & \left.=\left|C_{n n^{\prime}}\right|^{2}\left|\left\langle n^{0} \mathcal{N}\right|\left(\overline{n^{0}} n\right)(\bar{q} q)\right| n \mathcal{N}\right\rangle\left.\right|^{2} \\
& \sim\left|C_{n n^{\prime}} m_{\mathcal{N}}^{2} \sum_{q=u, d, s} f_{q}\right|^{2} \sigma_{n \mathcal{N}},
\end{aligned}
$$


[^0]:    In collaboration with IISER-TVM, University of Delhi, Punjab University, BHU, IIT Roorkee ....

